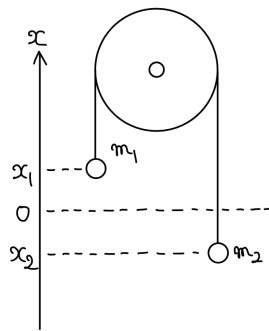


1 Lagrangian mechanics

1.1 Example: Atwood's machine



Invented by George Atwood in 1784 to verify the mechanical law of motion with constant acceleration.

- Constraint: $x_1 + x_2 = 0$
- 1 degree of freedom
- generalized coordinate: $x \equiv x_1 = -x_2$

$$T = \frac{1}{2}m_1\dot{x}_1^2 + \frac{1}{2}m_2\dot{x}_2^2 = \frac{1}{2}(m_1 + m_2)\dot{x}^2$$

$$V = m_1gx_1 + m_2gx_2 = (m_1 - m_2)gx$$

$$L = T - V = \frac{1}{2}(m_1 + m_2)\dot{x}^2 - (m_1 - m_2)gx$$

The Euler-Lagrange equation:

$$\frac{d}{dt} \frac{\partial L}{\partial \dot{x}} - \frac{\partial L}{\partial x} = (m_1 + m_2)\ddot{x} + (m_1 - m_2)g = 0$$

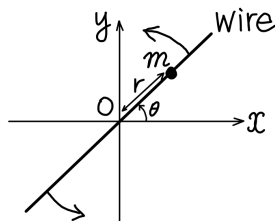
and thus

$$\ddot{x} = -\frac{m_1 - m_2}{m_1 + m_2}g$$

No constraint forces (tension) appear.

1.2 Example: A bead sliding on a uniformly rotating wire

An example of time-dependent constraints.



- Constraint: $\theta = \omega t$ (this is time-dependent!)
- 1 degree of freedom
- generalised coordinate: r

$$\begin{cases} x = r \cos \omega t \\ y = r \sin \omega t \end{cases}$$

Using the expression for T in polar coordinates,

$$T = \frac{1}{2}m(\dot{r}^2 + r^2\dot{\theta}^2) = \frac{1}{2}m(\dot{r}^2 + r^2\omega^2)$$

Note that this is not quadratic in \dot{r} .

If there are no other forces, $L = T$. The E-L equations are

$$\frac{d}{dt} \frac{\partial L}{\partial \dot{r}} - \frac{\partial L}{\partial r} = m\ddot{r} - mr\omega^2 = 0$$

$$\ddot{r} = r\omega^2$$

$$r = r_0 e^{\omega t}$$

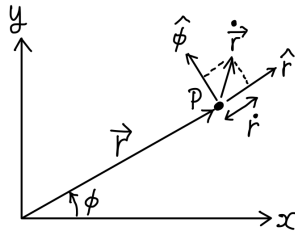
This means that the bead moves exponentially outward.

- Again no constraint force appears in the equation.
- The constraint force is *not* perpendicular to the motion, and thus it does work on the bead (energy is not conserved).

2 Curvilinear coordinates

Let us recall a few examples for curvilinear coordinate systems. These will be useful as generalised coordinates.

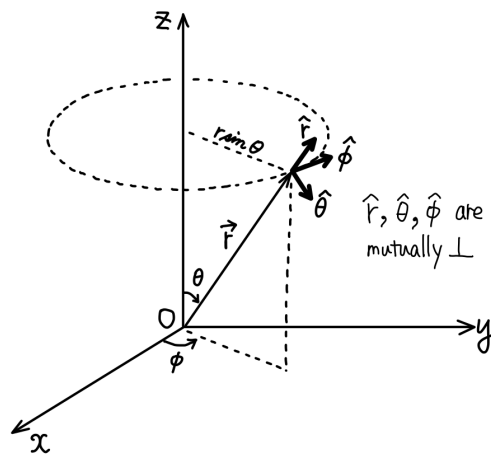
2.1 Polar coordinates



We have already seen this example in Lecture 9.

$$T = \frac{1}{2} m \dot{\vec{r}}^2 = \frac{1}{2} m (\dot{r}^2 + r^2 \dot{\phi}^2)$$

2.2 Spherical coordinates



$$\vec{r} = (x, y, z) = r(\sin \theta \cos \phi, \sin \theta \sin \phi, \cos \theta)$$

$$\theta \in [0, \pi] \quad : \quad \text{polar angle (colatitude)}$$

$$\phi \in [0, 2\pi] \quad : \quad \text{azimuthal angle}$$

- Components of the hatted vectors are

$$\hat{r} = (\sin \theta \cos \phi, \sin \theta \sin \phi, \cos \theta)$$

Take the θ derivative of this to get

$$\hat{\theta} = (\cos \theta \cos \phi, \cos \theta \sin \phi, -\sin \theta)$$

Finally,

$$\hat{\phi} = (-\sin \phi, \cos \phi, 0)$$

One can check that the vectors are orthonormal:

$$\hat{r}^2 = \hat{\theta}^2 = \hat{\phi}^2 = 1$$

$$\hat{r} \cdot \hat{\theta} = \hat{r} \cdot \hat{\phi} = \hat{\theta} \cdot \hat{\phi} = 0$$

and that $\hat{r} \cdot (\hat{\theta} \times \hat{\phi}) = +1$, i.e. they form a right-handed frame.

- In the kinetic term $\dot{\vec{r}}$ shows up which we need to compute

$$\dot{\vec{r}} = \frac{\partial \vec{r}}{\partial r} \dot{r} + \frac{\partial \vec{r}}{\partial \theta} \dot{\theta} + \frac{\partial \vec{r}}{\partial \phi} \dot{\phi}$$

We can find expressions for the partial derivatives using the hatted vectors

$$\frac{\partial \vec{r}}{\partial r} = \hat{r}, \quad \frac{\partial \vec{r}}{\partial \theta} = r \hat{\theta}, \quad \frac{\partial \vec{r}}{\partial \phi} = r \sin \theta \hat{\phi}$$

Plugging these back in gives

$$\dot{\vec{r}} = \dot{r} \hat{r} + r \dot{\theta} \hat{\theta} + r \sin \theta \dot{\phi} \hat{\phi}$$

and

$$T = \frac{m}{2} \dot{\vec{r}}^2 = \frac{m}{2} \left(\dot{r}^2 + r^2 (\dot{\theta}^2 + \sin^2 \theta \dot{\phi}^2) \right)$$

- The conjugate momenta are

$$p_r = \frac{\partial L}{\partial \dot{r}} = \frac{\partial T}{\partial \dot{r}} = m \dot{r}$$

$$p_\theta = \frac{\partial L}{\partial \dot{\theta}} = \frac{\partial T}{\partial \dot{\theta}} = m r^2 \dot{\theta}$$

$$p_\phi = \frac{\partial L}{\partial \dot{\phi}} = \frac{\partial T}{\partial \dot{\phi}} = m r^2 \sin^2 \theta \dot{\phi}$$

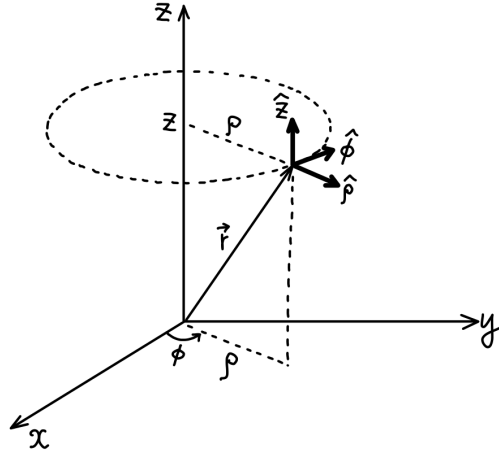
- Angular momentum (about origin)

$$\vec{L} = m \vec{r} \times \dot{\vec{r}} = m r \hat{r} \times (\dot{r} \hat{r} + r \dot{\theta} \hat{\theta} + r \sin \theta \dot{\phi} \hat{\phi}) = m r^2 (\underbrace{\dot{\theta} \hat{r} \times \hat{\theta}}_{\hat{\phi}} + \sin \theta \dot{\phi} \underbrace{\hat{r} \times \hat{\phi}}_{-\hat{\theta}}) = m r^2 \dot{\theta} \hat{\phi} - m r^2 \sin \theta \dot{\phi} \hat{\theta}$$

$$L_z = \vec{L} \cdot \hat{z} = m r^2 \dot{\theta} \underbrace{\hat{\phi} \cdot \hat{z}}_0 - m r^2 \sin \theta \dot{\phi} \underbrace{\hat{\theta} \cdot \hat{z}}_{-\sin \theta} = m r^2 \sin^2 \theta \dot{\phi}$$

So $L_z = p_\phi$ and it is conserved if ϕ is cyclic.

2.3 Cylindrical coordinates



$$\vec{r} = (\rho \cos \phi, \rho \sin \phi, z)$$

$$\dot{\vec{r}} = \frac{\partial \vec{r}}{\partial \rho} \dot{\rho} + \frac{\partial \vec{r}}{\partial \phi} \dot{\phi} + \frac{\partial \vec{r}}{\partial z} \dot{z} = \hat{\rho} \dot{\rho} + \rho \hat{\phi} \dot{\phi} + \hat{z} \dot{z}$$

where

$$\begin{aligned} \hat{r} &= \frac{\partial \vec{r}}{\partial \rho} = (\cos \phi, \sin \phi, 0) \\ \hat{\phi} &= \frac{1}{\rho} \frac{\partial \vec{r}}{\partial \phi} = (-\sin \phi, \cos \phi, 0) \\ \hat{z} &= \frac{\partial \vec{r}}{\partial z} = (0, 0, 1) \end{aligned}$$

These three vectors form a right-handed orthonormal basis.

$$T = \frac{1}{2} m \dot{\vec{r}}^2 = \frac{1}{2} m (\dot{x}^2 + \dot{y}^2 + \dot{z}^2) = \frac{1}{2} m (\dot{\rho}^2 + \rho^2 \dot{\phi}^2 + \dot{z}^2)$$

The conjugate momenta are

$$p_\rho = \frac{\partial L}{\partial \dot{\rho}} = m \dot{\rho}$$

$$p_\phi = \frac{\partial L}{\partial \dot{\phi}} = m \rho^2 \dot{\phi}$$

$$p_z = \frac{\partial L}{\partial \dot{z}} = m \dot{z}$$

$$\vec{L} = m \vec{r} \times \dot{\vec{r}} = m(\rho \hat{\rho} + z \hat{z}) \times (\dot{\rho} \hat{\rho} + \rho \dot{\phi} \hat{\phi} + \dot{z} \hat{z}) = m(\rho^2 \dot{\phi} \underbrace{\hat{\rho} \times \hat{\phi}}_{\hat{z}} + \rho \dot{z} \underbrace{\hat{\rho} \times \hat{z}}_{-\hat{\phi}} + z \dot{\rho} \underbrace{\hat{z} \times \hat{\rho}}_{\hat{\phi}} + z \rho \dot{\phi} \underbrace{\hat{z} \times \hat{\phi}}_{-\hat{\rho}})$$

$$\vec{L} = m \rho^2 \dot{\phi} \hat{z} + m(z \dot{\rho} - \rho \dot{z}) \hat{\phi} - m \rho z \dot{\phi} \hat{\rho}$$

$$L_z = \vec{L} \cdot \hat{z} = m \rho^2 \dot{\phi} = p_\phi$$

L_z is conserved if ϕ is cyclic.

Let us look at some examples of Lagrangian mechanics using these coordinate systems.

2.4 Example: A free particle in spherical coordinates

$$L = T = \frac{m}{2}(\dot{r}^2 + r^2(\dot{\theta}^2 + \sin^2 \theta \dot{\phi}^2))$$

The Euler-Lagrange equations

- for r :

$$\frac{d}{dt} \frac{\partial L}{\partial \dot{r}} = \frac{\partial L}{\partial r}$$

$$\frac{d}{dt}(m\dot{r}) = m\ddot{r} = mr(\dot{\theta}^2 + \sin^2 \theta \dot{\phi}^2)$$

- for θ :

$$\frac{d}{dt} \frac{\partial L}{\partial \dot{\theta}} = \frac{\partial L}{\partial \theta}$$

$$\frac{d}{dt}(mr^2\dot{\theta}) = mr^2 \sin \theta \cos \theta \dot{\phi}^2$$

- for ϕ :

$$\frac{d}{dt} \frac{\partial L}{\partial \dot{\phi}} = \frac{\partial L}{\partial \phi} = 0 \quad \text{cyclic coordinate}$$

$$\frac{d}{dt}(mr^2 \sin^2 \theta \dot{\phi}) = 0 \quad \Rightarrow \quad p_{\phi} = mr^2 \sin^2 \theta \dot{\phi} = \text{const.}$$

2.5 Example: A free particle in cylindrical coordinates

$$L = T = \frac{1}{2}m(\dot{\rho}^2 + \rho^2\dot{\phi}^2 + \dot{z}^2)$$

The Euler-Lagrange equations

- for z :

$$\frac{\partial L}{\partial z} = 0 \quad \text{cyclic coordinate}$$

thus the linear momentum along z is constant:

$$p_z = m\dot{z} = \text{const.}$$

- for ρ :

$$\frac{d}{dt} \frac{\partial L}{\partial \dot{\rho}} = \frac{\partial L}{\partial \rho}$$

$$m\ddot{\rho} = \underbrace{m\rho\dot{\phi}^2}_{\text{centrifugal force}}$$

- for ϕ :

$$\frac{d}{dt} \frac{\partial L}{\partial \dot{\phi}} = \frac{\partial L}{\partial \phi} = 0 \quad \text{cyclic coordinate}$$

$$p_{\phi} = m\rho^2\dot{\phi} = \text{const.} \equiv l$$

Plugging this into the E-L equation for ρ ,

$$m\ddot{\rho} = m\rho \left(\frac{l}{m\rho^2} \right)^2 = \frac{l^2}{m\rho^3}$$

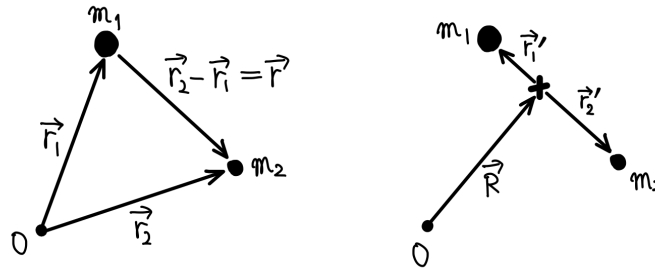
The RHS is the centrifugal force.

3 The Central Force Problem

As an important application of the Lagrangian formulation of mechanics, let us study the two-body problem with central forces.

3.1 Reduction to a one-body problem

Consider a system of two particles with masses m_1 and m_2 .



- $\vec{r}_1, \vec{r}_2 \Rightarrow$ there are 6 degrees of freedom
- As generalized coordinates, take:
 - COM position \vec{R}
 - difference vector $\vec{r} \equiv \vec{r}_2 - \vec{r}_1$

What is the Lagrangian?

Recall the decomposition of T in many-particle systems (see the end of Lecture 6):

$$T = \underbrace{\frac{1}{2}M\dot{\vec{R}}^2}_{\text{COM}} + \underbrace{\sum_{i=1}^N \frac{1}{2}m_i\dot{\vec{r}}_i'^2}_{\text{relative to COM}}$$

For two particles,

$$\begin{aligned} \vec{r}_1' &= \vec{r}_1 - \vec{R} = \vec{r}_1 - \frac{m_1\vec{r}_1 + m_2\vec{r}_2}{m_1 + m_2} = -\frac{m_2(\vec{r}_2 - \vec{r}_1)}{m_1 + m_2} = -\frac{m_2}{m_1 + m_2}\vec{r} \\ \vec{r}_2' &= \vec{r}_2 - \vec{R} = \vec{r}_2 - \frac{m_1\vec{r}_1 + m_2\vec{r}_2}{m_1 + m_2} = \frac{m_1(\vec{r}_2 - \vec{r}_1)}{m_1 + m_2} = \frac{m_1}{m_1 + m_2}\vec{r} \end{aligned}$$

Thus,

$$T = \frac{1}{2}M\dot{\vec{R}}^2 + \frac{1}{2}m_1 \left(-\frac{m_2}{m_1+m_2}\dot{\vec{r}} \right)^2 + \frac{1}{2}m_2 \left(\frac{m_1}{m_1+m_2}\dot{\vec{r}} \right)^2 = \frac{1}{2}M\dot{\vec{R}}^2 + \frac{1}{2}\frac{m_1m_2}{m_1+m_2}\dot{\vec{r}}^2$$

and we get

$$T = \frac{1}{2}M\dot{\vec{R}}^2 + \frac{1}{2}\mu\dot{\vec{r}}^2$$

where $\mu \equiv \frac{m_1m_2}{m_1+m_2}$ is the **reduced mass**.

Assume that the potential depends only on the relative position: $V = V(\vec{r})$. Then,

$$L_{\text{total}} = \underbrace{\frac{1}{2}M\dot{\vec{R}}^2}_{L_{\text{COM}}(\dot{\vec{R}})} + \underbrace{\frac{1}{2}\mu\dot{\vec{r}}^2 - V(\vec{r})}_{L_{\text{rel}}(\vec{r},\dot{\vec{r}})}$$

Note that the COM motion and the relative motion *decouple* from each other.

Explicitly,

- E-L equation for \vec{R} :

$$\frac{d}{dt} \frac{\partial L}{\partial \dot{\vec{R}}} = \frac{\partial L}{\partial \vec{R}} = 0 \quad \Rightarrow \quad M\ddot{\vec{R}} = 0 : \quad \text{trivial inertial motion}$$

- E-L equation for \vec{r} :

$$\frac{d}{dt} \frac{\partial L}{\partial \dot{\vec{r}}} = \frac{\partial L}{\partial \vec{r}} = 0 \quad \Rightarrow \quad \mu\ddot{\vec{r}} = -\frac{\partial}{\partial \vec{r}}V(\vec{r}) : \quad \text{motion of a particle with Lagrangian } L_{\text{rel}}$$

The two-body problem thus reduces to a one-body problem. The #DoF has been reduced to 3.

3.2 Central force

Now consider just the relative motion described by \vec{r} .

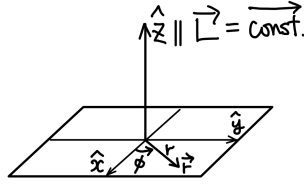
Assume $V = V(r)$ where $r \equiv |\vec{r}|$.

- The force is central: $\vec{F} \parallel \vec{r}$

$$\vec{F} = -\frac{\partial V}{\partial \vec{r}} = -V'(r)\hat{r}$$

- Angular momentum \vec{L} is conserved
- Motion is planar (since \vec{r} is in a plane perpendicular to \vec{L}). Thus, the #DoF is reduced to 2.

Take polar coordinates (r, ϕ) on the plane of motion.



$$L = T - V = \frac{1}{2}\mu(\dot{r}^2 + r^2\dot{\phi}^2) - V(r)$$

The Euler-Lagrange equation for ϕ :

$$\underbrace{\frac{d}{dt} \frac{\partial L}{\partial \dot{\phi}}}_{p_{\phi}} = \frac{\partial L}{\partial \phi} = 0 \quad \Rightarrow \quad \dot{p}_{\phi} = 0$$

$$p_{\phi} = \mu r^2 \dot{\phi} = \text{const.} \equiv l \quad (1)$$

Claim: $p_{\phi} = L_z$.

Proof:

$$\vec{r} = (r \cos \phi, r \sin \phi, 0)$$

$$\dot{\vec{r}} = (\dot{r} \cos \phi - r \sin \phi \dot{\phi}, \dot{r} \sin \phi + r \cos \phi \dot{\phi}, 0)$$

Let's take the cross-product,

$$\vec{r} \times \dot{\vec{r}} = (0, 0, r^2 \dot{\phi})$$

$$\vec{L} = \vec{r} \times \mu \dot{\vec{r}} = (0, 0, \mu r^2 \dot{\phi})$$

so indeed $L_z = p_{\phi}$.

3.3 Reduction to one-dimensional problem

Energy is conserved,

$$E = T + V = \frac{\mu}{2}(\dot{r}^2 + r^2\dot{\phi}^2) + V(r)$$

using $\dot{\phi} = \frac{l}{\mu r^2}$ from eqn. (1),

$$E = \frac{1}{2}\mu\dot{r}^2 + \frac{1}{2}\mu r^2 \left(\frac{l}{\mu r^2} \right)^2 + V(r) = \frac{1}{2}\mu\dot{r}^2 + \frac{l^2}{2\mu r^2} + V(r)$$

or

$$E = \frac{1}{2}\mu\dot{r}^2 + V_{\text{eff}}(r) \quad V_{\text{eff}}(r) = V(r) + \frac{l^2}{2\mu r^2}$$

V_{eff} is the **effective potential**.

- This is the expression for the energy of a particle in one dimension with potential V_{eff} .
- The 2d problem has been reduced to a 1d problem (#DoF: $6 \rightarrow 3 \rightarrow 2 \rightarrow 1$)
- The extra “force” is

$$-\frac{d}{dr} \left(\frac{l^2}{2\mu r^2} \right) = \frac{l^2}{\mu r^3}$$

This is nothing but the centrifugal force

$$F_{\text{cf}} = \frac{\mu v^2}{r}, \quad v = r\dot{\phi} = \frac{l}{\mu r} \quad \Rightarrow \quad F_{\text{cf}} = \frac{l^2}{\mu r^3}$$

The Newton equation is

$$\mu\ddot{r} = \frac{l^2}{\mu r^3} - V'(r)$$

Note: It would have been incorrect to replace $\dot{\phi}$ in the Lagrangian by $\dot{\phi} = \frac{l}{\mu r^2}$ as we did in the formula for the energy. This would lead to a wrong effective potential $V_{\text{eff}}^{\text{WRONG}}(r) = V(r) - \frac{l^2}{2\mu r^2}$ which has the wrong sign for the second term. This is because the Lagrangian formulation assumes that the dynamical variables are independent (see p. 140 of Hand & Finch).